# Identification of the dilute regime in particle sedimentation 

Pierre-Emmanuel Jabin<br>Département de Mathématiques et Applications, Ecole Normale Supérieure<br>45 rue d'Ulm, 75230 Paris Cedex 05, France, Felix Otto<br>Institut für Angewandte Mathematik, Universität Bonn,<br>Wegelerstrasse 10, 53115 Bonn


#### Abstract

We investigate the dynamics of rigid, spherical particles of radius $R$ sinking in a viscous fluid. Both the inertia of the particles and the fluid are neglected. We are interested in a large number $N$ of particles with average distance $d \gg R$. We investigate in which regime (in terms of $N$ and $R / d$ ) the particles do not significantly interact and approximately sink like single particles. We rigorously establish the lower bound $N_{\text {crit }} \geq C\left(\frac{d}{R}\right)^{3 / 2}$ for the critical number $N_{\text {crit }}$ of particles. This lower bound agrees with the heuristically expected $N_{\text {crit }}$ in terms of its scaling in $R / d$. The main difficulty lies in showing that the particles cannot get significantly closer over a relevant time scale. We use the method of reflection for the Stokes operator to bound the strength of the particle interaction.


## 1 Introduction

### 1.1 Motivation of the result

We consider the sedimentation of rigid spherical particles of the same radius in a fluid. The particles interact through the fluid: When one particle moves, it generates a fluid flow which acts on all the other particles.

We neglect the inertia of both particles and fluid. In particular, the fluid flow is quasi stationary and described by the incompressible Stokes system with no-slip boundary condition at the particles' surface. Hence the dynamics are driven by $\rho$, the difference in weight density between the particles and the fluid, and $e$, the gravity field. They are limited by the viscosity $\mu$ of the fluid. The quasi stationarity assumption on the fluid flow means that the fluid immediately adapts itself to the situation created by the particles' positions and their velocities: It does not "remember" anything about the past of the dynamics. We also neglect the rotation of the particles.

We always assume that the Stokes fluid is at rest at infinity, meaning that the fluid velocity $u$ satisfies $u \rightarrow 0$ for $|x| \rightarrow \infty$. In such an environment, a single particle of radius $R$ sinks with a velocity

$$
\begin{equation*}
V_{\text {single }}:=\frac{\rho}{6 \pi \mu} R^{2}|e| . \tag{1.1}
\end{equation*}
$$

A small number $N$ of distant particles will only show little interaction; they will sink like a single particle. Hence the velocity $V_{\text {cloud }}$ of such a cloud of particles is approximately equal to $V_{\text {single }}$ :

$$
V_{\text {cloud }} \approx V_{\text {single }}
$$

We call this the "non-interacting scenario". But if the number $N$ of particles is large, their interaction may no longer be a small perturbation. In such a regime, the fact that the fluid at rest at infinity (that is, $u \rightarrow 0$ for $|x| \rightarrow \infty$ ) may be "screened" from the particles in the interior of the cloud. Hence it is plausible that a macroscopic fluid flow is generated, which makes the cloud sink faster. We call this the "screening scenario". Our goal is to identify the cross-over between these two regimes. The critical number of particles $N_{\text {crit }}$ will depend on their radius $R$ and their average distance $d$, more precisely, on the non-dimensional ratio $R / d$. In this paper, we investigate the scaling of $N_{\text {crit }}$ in $R / d$. The main result is a rigorous lower bound for $N_{\text {crit }}$ as a function of $R / d$, see Theorem 1.1. We now will give a heuristic argument that this bound is optimal in terms of scaling.

Let us perform a little Gedankenexperiment: The screening scenario is mimicked by treating the cloud of particles as a single "meta-particle". This meta-particle would have diameter $\tilde{R} \sim N^{1 / 3} d$ and a difference $\tilde{\rho} \sim \rho\left(\frac{R}{d}\right)^{3}$ in density with respect to the fluid. Hence it would sink with a velocity which
scales as

$$
\begin{equation*}
\left|\tilde{V}_{\text {cloud }}\right| \sim \frac{\tilde{\rho}}{\mu} \tilde{R}^{2}|e| \sim \frac{\rho}{\mu} N^{2 / 3} \frac{R^{3}}{d}|e| . \tag{1.2}
\end{equation*}
$$

Folklore suggests that the screening scenario should prevail over the noninteracting scenario, if and only if the velocity $\tilde{V}_{\text {cloud }}$ is larger than the velocity $V_{\text {cloud }}$. Hence one obtains from (1.1) and (1.2) the following guess for the cross-over number $N_{\text {crit }}$

$$
\begin{equation*}
N_{\text {crit }}:=\left(\frac{d}{R}\right)^{3 / 2} \tag{1.3}
\end{equation*}
$$

Unfortunately, by this logic, we would have to choose yet another scenario, the "clustering scenario", where the particles form (possibly transient) clusters within the cloud. The extreme case of the clustering scenario is mimicked by a single "mega-particle". It would have diameter $R_{*} \sim N^{1 / 3} R$ and hence sink with a velocity which scales as

$$
\left|V_{*}\right| \sim \frac{\rho}{\mu} R_{*}{ }^{2}|e| \sim \frac{\rho}{\mu} N^{2 / 3} R^{2}|e|
$$

beating the velocity of the non-interacting scenario $V_{\text {cloud }}$. Nonetheless, we will show in this paper that as long as

$$
N \ll N_{\text {crit }},
$$

with $N_{\text {crit }}$ defined as in (1.3), the particles do not interact and all sink with approximately the velocity $V_{\text {single }}$ of a single particle.
In view of the clustering scenario, it is not surprising that the main issue is to rule out that particles get too close to each other. Let us make this more precise and denote by $d_{\text {min }}(t)$ the minimal distance between the particles at time $t$. We will assume that initially, all particles are well-separated in the sense of

$$
R \ll d_{\min }(0)
$$

The main effort consists in establishing that no clustering occurs over the relevant time scale

$$
R \ll d_{\text {min }}(t) \text { for } t=O(\tau)
$$

What is the relevant time scale? It is the time $\tau$ it takes for a single particle to sink a distance $N^{1 / 3} d$ of the order of the cloud diameter. More precisely, we set

$$
\begin{equation*}
\tau:=\frac{N^{1 / 3} d_{\min }(0)}{\left|V_{\text {single }}\right|} \tag{1.4}
\end{equation*}
$$

We also denote by $\Lambda$ the critical parameter

$$
\begin{equation*}
\Lambda(t):=\frac{R N^{2 / 3}}{d_{\min }(t)} \tag{1.5}
\end{equation*}
$$

The main result of this paper is
Theorem 1.1 There exists a universal constant $C<\infty$ such that if

$$
C R \leq d_{\min }(0), \quad N \geq C, \quad \text { and } \quad C \Lambda(0) \leq 1
$$

we have, where $V_{k}$ denotes the velocity of particle $k$,

$$
\begin{aligned}
d_{\min }(t) & \geq \frac{1}{2} d_{\min }(0), \\
\sup _{k}\left|V_{k}(t)-V_{\text {single }}\right| & \leq C \Lambda(0)\left|V_{\text {single }}\right|
\end{aligned}
$$

for all $t$ with

$$
0 \leq t \leq C^{-1} \Lambda^{-2}(0) \tau
$$

Theorem 1.1 shows that the non-interacting scenario prevails for $\Lambda \ll 1$ as heuristically expected.

### 1.2 The underlying dynamics

Let us now give the precise equations of the dynamics. Let $i=1, \cdots, N$ be an enumeration of the particles. Each particle is assumed to be a spherical ball of radius $R$ and center $X_{i}$, moving with the velocity $V_{i}$; we denote the ball by $B_{i}$.
At a given time $t$, the fluid flow is described by its pressure field $p \in \mathbb{R}$, its velocity field $u \in \mathbb{R}^{3}$ and its stress field $\sigma \in \operatorname{Sym}\left(\mathbb{R}^{3}\right)$ related by

$$
\begin{equation*}
\sigma_{\alpha, \beta}=-p \delta_{\alpha, \beta}+\mu\left(\partial_{\alpha} u_{\beta}+\partial_{\beta} u_{\alpha}\right) . \tag{1.6}
\end{equation*}
$$

The flow is assumed to be incompressible, to be at rest at infinity:

$$
\begin{equation*}
\nabla \cdot u=0 \quad \text { in } \mathbb{R}^{3}-\bigcup_{i} B_{i} \quad \text { and } \quad u \rightarrow 0 \quad \text { for }|x| \uparrow \infty \tag{1.7}
\end{equation*}
$$

and to satisfy the no-slip boundary condition at the particle interfaces

$$
\begin{equation*}
u=V_{i} \quad \text { on } \partial B_{i} \quad 1 \leq i \leq N . \tag{1.8}
\end{equation*}
$$

Finally, we postulate quasi stationary balance between viscous and gravitational forces in the bulk and at the interface

$$
\begin{equation*}
\nabla \cdot \sigma=0 \quad \text { in } \mathbb{R}^{3}-\bigcup_{i} B_{i} \quad \text { and } \quad \int_{\partial B_{i}} \sigma \cdot n d S+\rho R^{3} e=0,1 \leq i \leq N \tag{1.9}
\end{equation*}
$$

where $n$ denotes the outer normal of $B_{i}$. Notice that $\nabla \cdot \sigma$ turns into the Stokes equation for $u$ and $p$.

The set of assumptions (1.6), (1.7), (1.9) \& (1.8) are supposed to be a good model of the real dynamics if the particles are small enough so that the Reynolds number of the fluid is much less than 1 but not too small so as to neglect Brownian motion, and if the density of the particles is comparable to the density of the fluid (so the inertia of the particles is of the same order as the one of the fluid and hence negligible too). This model is used for the computation of sedimentation velocity (see [1], [6] or [9] for instance).
These dynamics are mathematically well-behaved as long as particles do not get too close. In fact, B. Desjardins and M. Esteban have shown that the dynamics - even with Navier-Stokes replacing quasi stationary Stokes - have a solution until the time of first collision between particles (see [5] and [14] for numerical simulations in this case). Notice however that it is not known whether collisions actually occur in this model. Even worse, the assumptions on which the model is based are no longer physically valid in a spatial and temporal neighborhood of a collision. In [11], it is shown that in two space dimensions, a particle cannot "bump" into a boundary. This indicates that in two space dimensions, collisions must occur at zero velocity. This corresponds to the result of [18] where a weak solution of the dynamics is defined if there is only one particle. But of course avoiding collisions is not enough for the asymptotics in $N$.

For a large number of particles, solving (1.6), (1.7), (1.9) \& (1.8) becomes numerically prohibitive. It is therefore crucial to investigate whether certain macroscopic aspects of the particle distribution can be predicted without solving the underlying microscopic dynamics. For instance, one could aim at deriving an evolution of the macroscopic number density of the particles (in the physical space when the inertia of the particles is neglected and in the phase space when it is not). J. Rubinstein and J.B. Keller have thus studied a limit toward a macroscopic system for particles without inertia (see [16] and also [4] for a study of a limit system close to this one but with Brownian
motion). If the inertia of the particles is not neglected, it is possible to derive kinetic equations (see [13]). A modeling by kinetic equations has also been studied in the case of a perfect fluid with a potential flow (see [10] and [17]).

Unfortunately, all this limits are formal. In general, we know how to pass from the microscopic dynamics to a macroscopic evolution only when the interaction has no singularity. For kinetic equations, the Boltzmann-Grad limit has been proved by R. Illner and M. Pulvirenti in [12] (see also [3]) but only for a short time. Concerning Vlasov equations, the convergence is proved only for regularized forces (see the book by H. Spohn [19] for instance). This question is also connected with the convergence of particle methods (for the Vlasov-Poisson equation, more details can be found in [2], [20] and [21]).

However for macroscopic systems (when the velocity of one particle depends only on its position), the question appears to be easier; for instance, the limit of point vortex method to the Euler equations has been proved in dimension 2 in [7]. We see Theorem 1.1 as a first step towards rigorously deriving an evolution of the macroscopic number density. For instance, Theorem 1.1 identifies the meaning of "macroscopic": Sedimentation is self-averaging on length-scales up to $N_{\text {crit }}^{1 / 3} d=\frac{d^{3 / 2}}{R^{1 / 2}}$; hence the macroscopic number density should be defined with respect to this critical length-scale. Moreover, Theorem 1.1 gives us hope that particle collisions can be ruled out even in the limit of infinite interacting particles over the relevant time scales.

### 1.3 The structure of the dynamics

The system formed of (1.6), (1.7), (1.9) \& (1.8) defines an evolution of the particle centers $\left\{X_{i} \in \mathbb{R}^{3}\right\}_{i}$. The dynamics can be written as

$$
\dot{X}_{i}=V_{i}\left(X_{1}, \ldots, X_{N}\right), \quad 1 \leq i \leq N,
$$

where the velocity $V_{i}$ of each particle is only determined by the configuration at the same time, i.e all the positions $X_{j}$.
The equations (1.6), (1.7), (1.9) \& (1.8) are linear in the velocities $V_{i}$. Therefore, it is convenient to introduce the following tensor $\left\{g_{i j}\left(X_{1}, \ldots, X_{N}\right) \in\right.$ $\left.\operatorname{Sym}\left(\mathbb{R}^{3}\right)\right\}_{i, j}, 1 \leq i, j \leq N$; For any $U \in \mathbb{R}^{3}, g_{i j} U \in \mathbb{R}^{3}$ is defined through

$$
g_{i j} U=-\int_{\partial B_{i}} \sigma n
$$

where $(p, u, \sigma)$ denotes the solution of (1.6), (1.7) and

$$
\begin{gathered}
\nabla \cdot \sigma=0 \quad \text { in } \mathbb{R}^{3}-\bigcup_{k} B_{k} \\
u=U \quad \text { on } \partial B_{j} \quad \text { and } \quad u=0 \quad \text { on } \partial B_{k} \quad \text { for all } k \neq j .
\end{gathered}
$$

We observe that the tensor $\left\{g_{i j} \in \operatorname{Sym}\left(\mathbb{R}^{3}\right)\right\}_{i, j}$ is symmetric and positive definite on $\mathbb{R}^{3 N}$. Indeed, for $\left\{U_{i}^{1}\right\}_{i}$ and $\left\{U_{i}^{2}\right\}_{i}$ we have

$$
\sum_{1 \leq i, j \leq N} U_{i}^{1} \cdot g_{i j} U_{j}^{2}=2 \mu \int_{\mathbb{R}^{3}-\bigcup_{i} B_{i}} \sum_{1 \leq \alpha, \beta \leq 3} \partial_{\alpha} u_{\beta}^{1} \partial_{\alpha} u_{\beta}^{2} d x
$$

where $u^{1}$ and $u^{2}$ are the solutions of (1.6), (1.7), (1.9) \& (1.8) with $\left\{V_{i}\right\}_{i}$ replaced by $\left\{U_{i}^{1}\right\}_{i}$ resp. $\left\{U_{i}^{2}\right\}_{i}$.
With this notation, the velocities $V_{i}$ are given by the solution of the linear system

$$
\sum_{j=1}^{N} g_{i j} V_{j}=F_{i}, \quad 1 \leq i \leq N
$$

where the right hand side vector $\left\{F_{i}\right\}_{i}$ is the constant gravity force on each particle

$$
F_{i}:=\rho R^{3} e, \quad 1 \leq i \leq N .
$$

Notice that the dynamics have the structure of a gradient flow with respect to the metric tensor $\left\{g_{i j}\right\}_{i, j}$ and the gravity potential. It is clear that one has to control the metric tensor $\left\{g_{i j}\right\}_{i, j}$ as a function of the particle positions $\left\{X_{i}\right\}_{i}$ in order to control the dynamics. The control required is summarized in Proposition 1.1. We recall the notation

$$
d_{i j}:=\left|X_{i}-X_{j}\right|, \quad d_{\min }=\min _{i \neq j} d_{i j} \quad \text { and } \quad \lambda=\frac{R N^{1 / 3}}{d_{\min }} .
$$

In order to avoid writing generic universal constants, we from now on use the notation $\lesssim$ and $\ll$.

Proposition 1.1 Provided

$$
R \ll d_{\min } \quad \text { and } \quad \Lambda \ll 1
$$

we have

$$
\begin{align*}
\left|g_{11}-6 \pi \mu R \mathrm{id}\right| & \lesssim \mu R \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}  \tag{1.10}\\
\left|g_{12}\right| & \lesssim \mu R \frac{R}{d_{12}},  \tag{1.11}\\
\left|g_{13}-g_{23}\right| & \lesssim \mu R d_{12}\left(\left(\frac{R}{d_{13}^{2}}+\frac{R}{d_{23}^{2}}\right)+\left(\frac{1}{d_{13}}+\frac{1}{d_{23}}\right) \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}\right) . \tag{1.12}
\end{align*}
$$

## 2 From Proposition 1.1 to Theorem 1.1

We first give three lemmas which enable us to control the velocity of each particle and then we prove Theorem 1.1.

### 2.1 Control on the particles' velocities

We begin with a completely technical lemma which is also used in the last section of the paper

Lemma 2.1 For any exponent $k \in[0,2]$, we have

$$
\begin{equation*}
\sum_{j \neq i} \frac{1}{d_{i j}^{k}} \lesssim \frac{N^{1-k / 3}}{d_{\min }^{k}} \tag{2.1}
\end{equation*}
$$

Lemma 2.2 Provided $R \ll d_{\text {min }}$ and $\Lambda \ll 1$ we have

$$
\begin{equation*}
\sup _{k}\left|V_{k}-V_{\text {single }}\right| \lesssim \Lambda\left|V_{\text {single }}\right| . \tag{2.2}
\end{equation*}
$$

Lemma 2.3 Provided $R \ll d_{\min }$ and $\Lambda \ll 1$ we have for all $i \neq j$

$$
\begin{equation*}
\left|V_{i}-V_{j}\right| \lesssim \frac{R N^{1 / 3}}{d_{\min }^{2}} \sup _{k}\left|V_{k}\right| d_{i j} \tag{2.3}
\end{equation*}
$$

Proof of Lemma 2.1. Since the property we want to prove is invariant by translation, we may suppose that $i=1$ and $X_{1}=0$ without any loss of generality. After that, we rearrange the indices such that the sequence $\left|X_{1}\right|, \ldots,\left|X_{N}\right|$ is non decreasing.

Since, for any $i \neq j,\left|X_{i}-X_{j}\right| \geq d$, the balls $B\left(X_{i}, d_{\text {min }} / 2\right)$ and $B\left(X_{j}, d_{\text {min }} / 2\right)$ do not intersect. Moreover for any $i$ between 2 and $n$, we know that

$$
\bigcup_{j=1}^{i} B\left(X_{j}, d_{\min } / 2\right) \subset B\left(0, d_{\min } / 2+\sup _{j=1, ., i}\left|X_{j}\right|\right) \subset B\left(0,2\left|X_{i}\right|\right)
$$

Comparing the two volumes, we find that

$$
i\left(\frac{d_{\min }}{2}\right)^{3} \leq\left(2\left|X_{i}\right|\right)^{3}
$$

or

$$
\left|X_{i}\right| \geq \frac{1}{4} d_{m i n} i^{1 / 3}
$$

We finish the proof with an easy comparison of the sum with an integral

$$
\begin{aligned}
\sum_{i=2}^{N} \frac{1}{\left|X_{1}-X_{i}\right|^{k}} & \leq \frac{4^{k}}{d_{\min }^{k}} \sum_{i=2}^{N} i^{-k / 3} \leq \frac{4^{k}}{d_{\min }^{k}} \int_{0}^{N} x^{-k / 3} d x \\
& \leq \frac{4^{k}}{d_{\min }^{k}}\left[\frac{X^{1-k / 3}}{1-k / 3}\right]_{0}^{N} \leq \frac{4^{k}}{1-k / 3} \times \frac{N^{1-k / 3}}{d_{\min }^{k}} \leq 48 \frac{N^{1-k / 3}}{d_{\min }^{k}}
\end{aligned}
$$

Proof of Lemma 2.2. We start by observing that for any $i$, we have on one hand

$$
\sum_{k} g_{i k} V_{k}-6 \pi \mu R V_{\text {single }}=F_{i}-\rho R^{3} e=0
$$

and on the other hand

$$
\begin{aligned}
\sum_{k} g_{i k} V_{k}-6 \pi \mu R V_{\text {single }} & =(6 \pi \mu R)\left(V_{i}-V_{\text {single }}\right) \\
& +\left(g_{i i}-6 \pi \mu R\right) V_{i} \\
& +\sum_{k \neq i} g_{i k} V_{k} .
\end{aligned}
$$

Therefore,

$$
\begin{equation*}
6 \pi \mu R\left|V_{i}-V_{\text {single }}\right| \leq\left|g_{i i}-6 \pi \mu R \mathrm{id}\right| \times \sup _{k}\left|V_{k}\right|+\sum_{k \neq i}\left|g_{i k}\right| \times \sup _{k}\left|V_{k}\right| . \tag{2.4}
\end{equation*}
$$

According to Prop. 1.1 we have

$$
\begin{aligned}
\left|g_{i i}-6 \pi \mu R \mathrm{id}\right| & \lesssim \mu R \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}, \\
\sum_{k \neq i}\left|g_{i k}\right| & \lesssim \mu R^{2} \sum_{k \neq i} \frac{1}{d_{i k}},
\end{aligned}
$$

which yields with Lemma 2.1

$$
\sum_{k \neq i}\left|g_{i k}\right| \lesssim \mu R \frac{R N^{2 / 3}}{d_{\min }} \lesssim \mu R \Lambda .
$$

We may consequently deduce from (2.4)

$$
6 \pi \mu R\left|V_{i}-V_{\text {single }}\right| \lesssim \mu R\left(\frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}+\Lambda\right) \sup _{k}\left|V_{k}\right| .
$$

Note also that

$$
\frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}=\frac{R N^{-1 / 3}}{d_{\min }} \times \frac{R N^{2 / 3}}{d_{\min }} \leq \frac{R}{d_{\min }} \Lambda \leq \Lambda,
$$

so eventually,
$6 \pi \mu R\left|V_{i}-V_{\text {single }}\right| \lesssim \mu R \Lambda \sup _{k}\left|V_{k}\right| \lesssim \mu R \Lambda \sup _{k}\left|V_{k}-V_{\text {single }}\right|+\mu R \Lambda V_{\text {single }}$.
Taking the supremum in $i$, we obtain

$$
\sup _{i}\left|V_{i}-V_{\text {single }}\right| \lesssim \Lambda \sup _{i}\left|V_{i}-V_{\text {single }}\right|+\Lambda\left|V_{\text {single }}\right| .
$$

This yields Lemma 2.2 since $\Lambda \ll 1$.

Proof of Lemma 2.3. The argument is almost the same; For any $i \neq j$, we write

$$
\sum_{k} g_{i k} V_{k}-\sum_{k} g_{j k} V_{k}=F_{i}-F_{j}=0
$$

whereas we also have

$$
\begin{aligned}
\sum_{k} g_{i k} V_{k}-\sum_{k} g_{j k} V_{k} & =(6 \pi \mu R)\left(V_{i}-V_{j}\right) \\
& +\left(g_{i i}-6 \pi \mu R\right) V_{i}-\left(g_{j j}-6 \pi \mu R\right) V_{j} \\
& +g_{i j} V_{j}-g_{j i} V_{i} \\
& +\sum_{k \neq i, k \neq j}\left(g_{i k}-g_{j k}\right) V_{k}
\end{aligned}
$$

Thus, as in the previous lemma

$$
\begin{align*}
6 \pi \mu R\left|V_{i}-V_{j}\right| & \leq\left\{2 \sup _{k}\left|g_{k k}-6 \pi \mu R \mathrm{id}\right|+2\left|g_{i j}\right|\right. \\
& \left.+\sum_{k \neq i, k \neq j}\left|g_{i k}-g_{j k}\right|\right\} \sup _{k}\left|V_{k}\right| \tag{2.5}
\end{align*}
$$

We apply Proposition 1.1:

$$
\begin{aligned}
\sup _{k}\left|g_{k k}-6 \pi \mu R \mathrm{id}\right| & \lesssim \mu R \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}} \leq \mu R \frac{R}{d_{\min }} \frac{R N^{2 / 3}}{d_{\min }}=\mu R \frac{R}{d_{\min }} \Lambda \\
\left|g_{i j}\right| \lesssim & \lesssim \mu \frac{R}{d_{i j}} \lesssim \mu R \frac{R}{d_{\min }} \\
\sum_{k \neq i, k \neq j}\left|g_{i k}-g_{j k}\right| \lesssim & \mu R\left\{d_{i j} R\left(\sum_{k \neq i} \frac{1}{d_{i k}^{2}}+\sum_{k \neq j} \frac{1}{d_{j k}^{2}}\right)\right. \\
& \left.+d_{i j} \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}\left(\sum_{k \neq i} \frac{1}{d_{i k}}+\sum_{k \neq j} \frac{1}{d_{j k}}\right)\right\}
\end{aligned}
$$

According to Lemma 2.1, the last estimate turns into

$$
\begin{aligned}
\sum_{k \neq i, k \neq j}\left|g_{i k}-g_{j k}\right| & \lesssim \mu R d_{i j}\left\{\frac{R N^{1 / 3}}{d_{\min }^{2}}+\frac{R^{2} N}{d_{\min }^{3}}\right\} \\
& =\mu R d_{i j}\left\{\frac{R N^{1 / 3}}{d_{\min }^{2}}+\frac{R N^{1 / 3}}{d_{\min }^{2}} \times \frac{R N^{2 / 3}}{d_{\min }}\right\} \\
& =\mu R d_{i j} \frac{R N^{1 / 3}}{d_{\min }^{2}}(1+\Lambda)
\end{aligned}
$$

Inserting these three estimates into (2.5) yields

$$
\left|V_{i}-V_{j}\right| \lesssim\left\{\frac{R}{d_{\min }}(1+\Lambda)+d_{i j} \frac{R N^{1 / 3}}{d_{\min }^{2}}(1+\Lambda)\right\} \sup _{k}\left|V_{k}\right|,
$$

and we conclude by observing that $\Lambda \ll 1$ and

$$
d_{i j} \frac{R N^{1 / 3}}{d_{\min }^{2}} \geq \frac{R N^{1 / 3}}{d_{\min }} \geq \frac{R}{d_{\min }} .
$$

### 2.2 Proof of Theorem 1.1

Let $C<\infty$ denote the maximum generic constant in the estimates of Lemmas 2.2 and 2.3. Let $t_{*}$ be such that

$$
d_{\min }(t) \geq \frac{1}{2} d_{\min }(0) \quad \text { for } 0 \leq t \leq t_{*}
$$

Then we have $\Lambda(t) \leq 2 \Lambda(0)$ for $o \leq t \leq t_{*}$ and therefore, provided $\Lambda(0)$ was small enough, we may apply all previous results up to time $t_{*}$.
According to Lemma 2.2, we have

$$
\sup _{k}\left|V_{k}(t)-V_{\text {single }}\right| \lesssim \Lambda(0)\left|V_{\text {single }}\right|,
$$

and hence in particular, using $\Lambda(0) \ll 1$,

$$
\sup _{k}\left|V_{k}(t)\right| \lesssim\left|V_{\text {single }}\right|,
$$

Therefore, we obtain from Lemma 2.3

$$
\left|V_{i}(t)-V_{j}(t)\right| \lesssim \frac{R^{2} N}{d_{\min }(0)^{3}}\left|V_{\text {single }}\right| d_{i j}(t)
$$

Let $C$ denote the universal constant implicit in the above estimate, i. e.

$$
\left|V_{i}(t)-V_{j}(t)\right| \leq C \frac{R^{2} N}{d_{\min (0)^{3}}}\left|V_{\text {single }}\right| d_{i j}(t),
$$

Since

$$
\frac{d}{d t} d_{i j}(t) \geq-\left|V_{i}(t)-V_{j}(t)\right|
$$

this entails

$$
d_{i j}(t) \geq \exp \left(-C \frac{R^{2} N}{d_{\min }(0)^{3}}\left|V_{\text {single }}\right| t\right) d_{i j}(0)
$$

or

$$
d_{\min }(t) \geq \exp \left(-C \frac{R^{2} N}{d_{\min }(0)^{3}}\left|V_{\text {single }}\right| t\right) d_{\min }(0)
$$

for $0 \leq t \leq t_{*}$. We conclude by observing that

$$
\exp \left(-C \frac{R^{2} N}{d_{\min }(0)^{3}}\left|V_{\text {single }}\right| t_{*}\right) \geq \frac{1}{2}
$$

is ensured by

$$
C \frac{R^{2} N}{d_{\min }(0)^{3}}\left|V_{\text {single }}\right| t_{*} \leq 1 .
$$

On the other hand, by definition of $\tau$,

$$
\frac{R^{2} N}{d_{\text {min }}(0)^{3}}\left|V_{\text {single }}\right|=\Lambda^{2}(0) \frac{1}{\tau} .
$$

This shows that we can choose $t_{*}$ at least of order $\Lambda^{-2}(0) \tau$.

## 3 Proof of Proposition 1.1

We begin with a technical lemma, then we present the method we use to prove Proposition 1.1.

### 3.1 The single-particle solution

The main idea is to express the solution operator for the multiple-particle Stokes problem in terms of the single-particle solution operator $T_{j}$

$$
T_{j}: L_{0}^{\infty}\left(\partial B_{j}\right) \rightarrow L^{\infty}\left(\mathbb{R}^{3}-B_{j}\right)
$$

The single-particle solution operator $T_{j}$ is defined as follows: If $U$ is a bounded velocity field on $\partial B_{j}$ with zero average normal component (this defines the space $\left.L_{0}^{\infty}\left(\partial B_{j}\right)\right)$, then $u=T_{j} U$ is the solution of the Stokes equations in $\mathbb{R}^{3}-B_{j}$ with Dirichlet boundary condition $U$ on $\partial B_{j}$. The Ansatz of writing the solution operator of an elliptic equation with a composite Dirichlet boundary in terms of the solution operators for the individual components
is known as the "method of reflections". It capitalizes on the fact that the solution operator for the individual component is well-understood. In our case, $T_{j}$ satisfies the following estimates, where $\|\cdot\|_{\partial B_{j}}$ denotes the $L^{\infty}$ norm on $\partial B_{j}$ :

Lemma 3.1 If $R \ll d_{\text {min }}$

$$
\begin{align*}
\left\|T_{j} U\right\|_{\partial B_{i}} & \lesssim \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} \text { for } i \neq j,  \tag{3.1}\\
\left\|T_{j} U-\left(T_{j} U\right)\left(\cdot+X_{k}-X_{i}\right)\right\|_{\partial B_{i}} & \lesssim d_{i k}\left(\frac{R}{d_{i j}^{2}}+\frac{R}{d_{j k}^{2}}\right)\|U\|_{\partial B_{j}}, \text { for } i \neq j, i \neq k, \\
\left|\int_{\partial B_{j}} \sigma\left(T_{j} U\right) \cdot n d S\right| & \lesssim \mu R\|U\|_{\partial B_{j}} . \tag{3.2}
\end{align*}
$$

Proof of Lemma 3.1. This result is shown with the explicit integral formula for $u$ and $p$. Since the problem is symmetric by translation, we may suppose that $j=1$ with $X_{1}=0$. As it is explained in [15], we have

$$
\begin{aligned}
& u_{\alpha}(x)=\frac{1}{8 \pi \mu} \int_{\partial B_{1}} U_{\beta}(y)\left(\mu \frac{\partial T_{\beta \alpha}}{\partial n}-P_{\alpha} n_{\beta}\right) d S(y), \\
& p(x)=\frac{1}{4 \pi} \int_{\partial B_{1}} U_{\alpha}\left(\mu \frac{\partial S_{\alpha}}{\partial n}+q n_{\alpha}\right) d S(y) .
\end{aligned}
$$

Let us denote

$$
\begin{aligned}
& r_{0}=|x|, \quad \bar{x}=\frac{R^{2}}{r_{0}^{2}} x, \quad \bar{r}_{0}=|\bar{x}| \\
& z=y-x, \quad r=|z|, \quad \bar{z}=y-\bar{x}, \quad \bar{r}=|\bar{z}| .
\end{aligned}
$$

We can now give the formula for $T$

$$
\begin{aligned}
T_{\alpha \beta} & =\frac{\delta_{\alpha \beta}}{r}+\frac{z_{\alpha} z_{\beta}}{r^{3}}-\frac{R}{r_{0} \bar{r}} \delta_{\alpha \beta}-\frac{R^{3} \bar{z}_{\alpha} \bar{z}_{\beta}}{r_{0}^{3}} \overline{\bar{r}}^{3} \\
& -\frac{R}{r_{0}^{2}-R^{2}} \bar{r}_{0}^{3} \\
r_{0} & \left.\left.\frac{\bar{x}_{\alpha} \bar{x}_{\alpha}}{r^{3}}+x_{\beta} \bar{z}_{\alpha}\right)-2 \bar{x}_{\alpha} \frac{\bar{x}_{\alpha} \bar{x}_{\beta}}{R^{3}} \bar{x} \cdot \nabla_{y} \frac{1}{\bar{r}}\right)-\left(|y|^{2}-R^{2}\right) \frac{\partial \phi_{\beta}}{\partial y_{\alpha}},
\end{aligned}
$$

whereas for $P$ and $\phi$

$$
\begin{aligned}
& P_{\alpha}=2 \mu \frac{z_{\alpha}}{r^{3}}-2 \mu \frac{R^{3}}{r_{0}^{3}} \overline{\bar{z}^{3}}-2 \mu \phi_{\alpha}-4 \mu y \cdot \nabla_{y} \phi_{\alpha}, \\
& \phi_{\alpha}=\frac{r_{0}^{2}-R^{2}}{2 r_{0}^{3}}\left(\frac{3 x_{\alpha}}{R \bar{r}}+\frac{R \bar{z}_{\alpha}}{\bar{r}^{3}}+2 \frac{x_{\alpha}}{R} \bar{x} \cdot \nabla_{y} \frac{1}{\bar{r}}+3 \frac{R}{\bar{r}_{0}} \frac{\partial}{\partial \bar{x}_{\alpha}} \log \frac{\bar{r}_{0} \bar{r}+y \cdot \bar{x}-r_{0}^{2}}{|y| \bar{r}_{0}+y \cdot \bar{x}}\right) .
\end{aligned}
$$

And finally the formulas for $S$, and $q$ are

$$
\begin{aligned}
& S_{\alpha}=\frac{\partial}{\partial y_{\alpha}} \frac{1}{r}-\frac{R}{r_{0}^{3}} \frac{x}{\bar{r}}-\frac{R^{3}}{r_{0}^{3}} \frac{\partial}{\partial y_{\alpha}} \frac{1}{\bar{r}}+\frac{2 \bar{x}_{\alpha}}{R r_{0}} \bar{x} \cdot \nabla_{y} \frac{1}{\bar{r}}-\left(|y|^{2}-R^{2}\right) \frac{\partial \psi}{\partial y_{\alpha}}, \\
& q=2 \mu\left(\psi+2 y \cdot \nabla_{y} \psi\right),
\end{aligned}
$$

with

$$
\psi=\frac{1}{2 R r_{0}}\left(\frac{3}{\bar{r}}+2 \bar{x} \cdot \nabla_{y} \frac{1}{\bar{r}}+\frac{3}{\bar{r}_{0}} \log \frac{\bar{r}_{0} \bar{r}+y \cdot \bar{x}-r_{0}^{2}}{|y| \bar{r}_{0}+y \cdot \bar{x}}\right)
$$

With these expressions, it is easy (but lengthy) to check that the kernels of the two convolutions are quite regular away from $\partial B_{1}$, thus proving the lemma.

### 3.2 The multiple-particle solution

In order to express the multiple-particle solution operator in terms of the $T_{j}$ 's, it is convenient to introduce the operator

$$
A: \Pi_{j=1}^{N} L_{0}^{\infty}\left(\partial B_{j}\right) \longrightarrow \Pi_{j=1}^{N} L_{0}^{\infty}\left(\partial B_{j}\right)
$$

which is defined through

$$
A_{i j} U=\left\{\begin{array}{cl}
\left(T_{j}\left(U_{\mid \partial B_{j}}\right)\right)_{\mid \partial B_{i}} & \text { for } j \neq i  \tag{3.3}\\
U & \text { for } j=i
\end{array}\right\}
$$

The next lemma collects the estimates on $A$ and $A^{-1}$ we later need. We denote by $A_{i j}^{-1}$ the component $i, j$ of the inverse of $A$ (and not the inverse of $A_{i j}$ ).
Lemma 3.2 $A^{-1}$ exists for $\Lambda \ll 1, R \ll d_{\text {min }}$, and we have

$$
\begin{align*}
\left\|A_{i j} U\right\|_{\partial B_{i}} & \lesssim \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} \quad \text { for } j \neq i  \tag{3.4}\\
\left\|A_{i j}^{-1} U\right\|_{\partial B_{i}} & \lesssim \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} \quad \text { for } j \neq i  \tag{3.5}\\
\left\|A_{i i}^{-1} U-U\right\|_{\partial B_{i}} & \lesssim \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}\|U\|_{\partial B_{i}}  \tag{3.6}\\
\text { in particular }\left\|A_{i i}^{-1} U\right\|_{\partial B_{i}} & \lesssim\|U\|_{\partial B_{i}} \tag{3.7}
\end{align*}
$$

We now may express the multiple-particle solution operator and thus also the metric tensor $g_{i j}$ in terms of the single-particle solution operator $T_{i}$ (and the related $A_{i j}$ ). We recall that

$$
\begin{equation*}
\int_{\partial B_{j}} \sigma\left(T_{i} U\right) \cdot n d S=0 \quad \text { for } j \neq i \tag{3.8}
\end{equation*}
$$

and

$$
\begin{equation*}
\int_{\partial B_{i}} \sigma\left(T_{i} U\right) \cdot n d S=6 \pi \mu R U \quad \text { for constant } U \text {. } \tag{3.9}
\end{equation*}
$$

For a constant vector $U$, we mean by the notation $A_{i j} U$ or $A_{i j}^{-1} U$ the component $i$ of the operator $A$ or $A^{-1}$ applied to the function equal to $U$ on $\partial B_{j}$ and to 0 on any $\partial B_{k}$ with $k \neq j$.

Lemma 3.3 We have for any constant vector $U$

$$
g_{i j} U=\int_{\partial B_{i}} \sigma\left(T_{i} A_{i j}^{-1} U\right) \cdot n d S
$$

## Proof of Lemma 3.2.

Estimate (3.4) is an immediate consequence of the definition (3.3) and the estimate (3.1). We now argue that for all positive integers $n$

$$
\left\|(\mathrm{id}-A)_{i j}^{n} U\right\|_{\partial B_{i}} \leq\left\{\begin{array}{cl}
(C \Lambda)^{n-1} \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} & \text { for } j \neq i  \tag{3.10}\\
(C \Lambda)^{n-1} \frac{\Lambda}{N}\|U\|_{\partial B_{i}} & \text { for } j=i
\end{array}\right\}
$$

for some universal constant $C<\infty$. Again (id $-A)_{i j}^{n}$ is the component $i j$ of the operator id $-A$ to the power $n$ and not $i d-A_{i j}$ to the power $n$. To establish (3.10), we make the Ansatz

$$
\left\|(\mathrm{id}-A)_{i j}^{n} U\right\|_{\partial B_{i}} \leq\left\{\begin{array}{cl}
\alpha_{n} \Lambda^{n-1} \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} & \text { for } j \neq i  \tag{3.11}\\
\beta_{n} \Lambda^{n-1} \frac{\Lambda}{N}\|U\|_{\partial B_{i}} & \text { for } j=i
\end{array}\right\}
$$

and derive the condition

$$
\begin{align*}
& \alpha_{1} \geq C_{0} \text { and }  \tag{3.12}\\
& \beta_{1} \geq 0 \alpha_{n+1} \geq 2 C_{0} C_{1} \alpha_{n}+C_{0} \beta_{n}  \tag{3.13}\\
&
\end{align*}
$$

where $C_{0}, C_{1}$ and $C_{2}$ stand for the universal constants with

$$
\begin{align*}
\left\|A_{i j} U\right\|_{\partial B_{i}} & \leq C_{0} \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} \quad \text { for } j \neq i,  \tag{3.14}\\
\sum_{j \neq i} \frac{R}{d_{i j}} & \leq C_{1} \frac{R N^{2 / 3}}{d_{\min }}=C_{1} \Lambda,  \tag{3.15}\\
\sum_{j \neq i}\left(\frac{R}{d_{i j}}\right)^{2} & \leq C_{2} \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}=C_{2} \frac{\Lambda^{2}}{N} . \tag{3.16}
\end{align*}
$$

We establish (3.12) and (3.13) by induction in $n$. It is true for $n=1$ by (3.14). Assume now it holds for $n$. Because of

$$
(\mathrm{id}-A)_{i j}^{n+1} U=\sum_{k}(\mathrm{id}-A)_{i k}(\mathrm{id}-A)_{k j}^{n} U=-\sum_{k \neq i} A_{i k}(\mathrm{id}-A)_{k j}^{n} U,
$$

and (3.14), we have

$$
\begin{equation*}
\left\|(\mathrm{id}-A)_{i j}^{n+1} U\right\|_{\partial B_{i}} \leq \sum_{k \neq i} C_{0} \frac{R}{d_{i k}}\left\|(\mathrm{id}-A)_{k j}^{n} U\right\|_{\partial B_{k}} \tag{3.17}
\end{equation*}
$$

We first treat the case $j \neq i$; we obtain from the assumption (3.11)

$$
\begin{equation*}
\left\|(\mathrm{id}-A)_{i j}^{n+1} U\right\|_{\partial B_{i}} /\|U\|_{\partial B_{j}} \leq\left(\sum_{k \neq i, j} C_{0} \frac{R}{d_{i k}} \alpha_{n} \Lambda^{n-1} \frac{R}{d_{k j}}\right)+C_{0} \frac{R}{d_{i j}} \beta_{n} \Lambda^{n-1} \frac{\Lambda}{N} . \tag{3.18}
\end{equation*}
$$

We now use the triangle inequality $d_{i j} \leq d_{i k}+d_{k j}$ in form of $\frac{1}{d_{i k} d_{k j}} \leq \frac{1}{d_{i j}}\left(\frac{1}{d_{i k}}+\right.$ $\frac{1}{d_{k j}}$ ). Hence we have by (3.15)

$$
\sum_{k \neq i, j} \frac{R}{d_{i k}} \frac{R}{d_{k j}} \leq 2 C_{1} \Lambda \frac{R}{d_{i j}}
$$

Thus (3.18) turns into

$$
\begin{aligned}
\Lambda^{-n}\left\|(\mathrm{id}-A)_{i j}^{n+1} U\right\|_{\partial B_{i}} /\|U\|_{\partial B_{j}} & \leq\left(2 C_{0} C_{1} \alpha_{n}+C_{0} \frac{1}{N} \beta_{n}\right) \frac{R}{d_{i j}} \\
& \leq\left(2 C_{0} C_{1} \alpha_{n}+C_{0} \beta_{n}\right) \frac{R}{d_{i j}}
\end{aligned}
$$

This establishes (3.12). We now treat the case $j=i$. Using (3.17) for $j=i$ and the assumption (3.11), we obtain

$$
\left\|(\mathrm{id}-A)_{i i}^{n+1}\right\|_{\partial B_{i}} /\|U\|_{\partial B_{i}} \leq \sum_{k \neq i} C_{0} \frac{R}{d_{i k}} \alpha_{n} \Lambda^{n-1} \frac{R}{d_{i k}} .
$$

Using (3.16), this turns into

$$
\Lambda^{-n}\left\|(\mathrm{id}-A)_{i i}^{n+1} U\right\|_{\partial B_{i}} /\|U\|_{\partial B_{i}} \leq C_{0} C_{2} \alpha_{n} \frac{\Lambda}{N}
$$

which yields (3.13). This establishes (3.10).
We formally have $A^{-1}=\sum_{n=0}^{\infty}(\mathrm{id}-A)^{n}$. We now show that this series converges and that the estimates (3.5) and (3.6) hold. Indeed, for $j \neq i$ we formally have

$$
A_{i j}^{-1} U=\sum_{n=1}^{\infty}(\mathrm{id}-A)_{i j}^{n} U
$$

and thus according to (3.10)

$$
\left\|A_{i j}^{-1} U\right\|_{\partial B_{i}} \leq \sum_{n=1}^{\infty}(C \Lambda)^{n-1} \frac{R}{d_{i j}}\|U\|_{\partial B_{j}} \lesssim \frac{R}{d_{i j}}\|U\|_{\partial B_{j}}
$$

For $j=i$ on the other hand we formally have

$$
A_{i i}^{-1} U=U+\sum_{n=2}^{\infty}(\mathrm{id}-A)_{i i}^{n} U
$$

and thus according to (3.10)

$$
\left\|A_{i i}^{-1} U-U\right\|_{\partial B_{i}} \leq \sum_{n=2}^{\infty}(C \Lambda)^{n-1} \frac{\Lambda}{N}\|U\|_{\partial B_{i}} \lesssim \frac{\Lambda^{2}}{N}\|U\|_{\partial B_{i}}=\frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}\|U\|_{\partial B_{i}}
$$

## Proof of Lemma 3.3.

Fix a particle $j$ and a constant velocity field $U$ on $\partial B_{j}$. Let $u$ denote the solution of the Stokes equations in $\mathbb{R}^{3}-\bigcup_{k} B_{k}$ with Dirichlet boundary data $\delta_{k j} U$ on $\partial B_{k}$. We claim that this solution can be written as

$$
\begin{equation*}
u=\sum_{\ell} T_{\ell} A_{\ell j}^{-1} U \tag{3.19}
\end{equation*}
$$

Indeed, the r. h. s. of (3.19) is a solution of the Stokes equations in $\bigcap_{k}\left(\mathbb{R}^{3}-\right.$ $\left.B_{k}\right)=\mathbb{R}^{3}-\bigcup_{k} B_{k}$. Its Dirichlet boundary data are as desired given by

$$
u_{\mid \partial B_{k}}=\sum_{\ell}\left(T_{\ell} A_{\ell j}^{-1} U\right)_{\mid \partial B_{k}}=\sum_{\ell} A_{k \ell} A_{\ell j}^{-1} U=\delta_{k j} U .
$$

This establishes (3.19).

We now fix a particle $i$ and observe that by definition of the metric tensor and thanks to (3.19) and (3.8)
$g_{i j} U=\int_{\partial B_{i}} \sigma(u) \cdot n d S=\sum_{\ell} \int_{\partial B_{i}} \sigma\left(T_{\ell} A_{\ell j}^{-1} U\right) \cdot n d S=\int_{\partial B_{i}} \sigma\left(T_{i} A_{i j}^{-1} U\right) \cdot n d S$.

### 3.3 Proof of Proposition 1.1

Estimate (1.10) is straightforward: According to Lemma 3.3 and to (3.9) we have

$$
g_{11} U-6 \pi \mu R U=\int_{\partial B_{1}} \sigma\left(T_{1}\left(A_{11}^{-1} U-U\right)\right) \cdot n d S
$$

Because of Lemma 3.1 and Lemma 3.2, this yields the desired estimate:

$$
\left|g_{11} U-6 \pi \mu R U\right| \lesssim \mu R\left\|A_{11}^{-1} U-U\right\|_{\partial B_{1}} \lesssim \mu R \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}|U| .
$$

Also estimate (1.11) is easy: By Lemma 3.3, Lemma 3.1 and Lemma 3.2 we have

$$
\left|g_{12} U\right| \lesssim \mu R\left\|A_{12}^{-1} U\right\|_{\partial B_{1}} \lesssim \mu R \frac{R}{d_{12}}|U| .
$$

We now tackle (1.12). We first derive the representation

$$
\begin{align*}
g_{13} U-g_{23} U= & \int_{\partial B_{1}} \sigma\left(T_{1}\left(u-u\left(\cdot+X_{2}-X_{1}\right)\right)\right) \cdot n d S \\
& +\int_{\partial B_{1}} \sigma\left(T_{1} v_{1}\right) \cdot n d S-\int_{\partial B_{2}} \sigma\left(T_{2} v_{2}\right) \cdot n d S \tag{3.20}
\end{align*}
$$

where

$$
\begin{align*}
u & :=T_{3}\left(-U+\sum_{j \neq 3} A_{3 j} A_{j 3}^{-1} U\right)+\sum_{i \neq 1,2,3} T_{i} \sum_{j \neq i} A_{i j} A_{j 3}^{-1} U,  \tag{3.21}\\
v_{1} & :=T_{2} \sum_{j \neq 2} A_{2 j} A_{j 3}^{-1} U,  \tag{3.22}\\
v_{2} & :=T_{1} \sum_{j \neq 1} A_{1 j} A_{j 3}^{-1} U . \tag{3.23}
\end{align*}
$$

Indeed, according to Lemma 3.3, we have

$$
g_{13} U-g_{23} U=\int_{\partial B_{1}} \sigma\left(T_{1} A_{13}^{-1} U\right) \cdot n d S-\int_{\partial B_{2}} \sigma\left(T_{2} A_{23}^{-1} U\right) \cdot n d S .
$$

Using

$$
A^{-1}=2 \mathrm{id}-A+(\mathrm{id}-A)^{2} A^{-1}
$$

we rewrite

$$
\begin{aligned}
A_{13}^{-1} & =-A_{13}+\sum_{i \neq 1} \sum_{j \neq i} A_{1 i} A_{i j} A_{j 3}^{-1} \\
& =A_{13}\left(-\mathrm{id}+\sum_{j \neq 3} A_{3 j} A_{j 3}^{-1}\right)+\sum_{i \neq 1,2,3} A_{1 i} \sum_{j \neq i} A_{i j} A_{j 3}^{-1}+A_{12} \sum_{j \neq 2} A_{2 j} A_{j 3}^{-1} .
\end{aligned}
$$

In view of (3.21) \& (3.22), this means

$$
A_{13}^{-1} U=\left(u+v_{1}\right)_{\mid \partial B_{1}} .
$$

Likewise, we have $A_{23}^{-1} U=\left(u+v_{2}\right)_{\mid \partial B_{2}}$. Hence we obtain

$$
\begin{aligned}
g_{13} U-g_{23} U= & \int_{\partial B_{1}} \sigma\left(T_{1} u\right) \cdot n d S+\int_{\partial B_{1}} \sigma\left(T_{1} v_{1}\right) \cdot n d S \\
& -\int_{\partial B_{2}} \sigma\left(T_{2} u\right) \cdot n d S-\int_{\partial B_{2}} \sigma\left(T_{2} v_{2}\right) \cdot n d S .
\end{aligned}
$$

Because of $\int_{\partial B_{2}} \sigma\left(T_{2} u\right) \cdot n d S=\int_{\partial B_{1}} \sigma\left(T_{1}\left(u\left(\cdot+X_{2}-X_{1}\right)\right)\right) \cdot n d S$, this turns into (3.20).

We now estimate $g_{13} U-g_{23} U$ based on the representation (3.20). Using Lemma 3.1, we have

$$
\begin{equation*}
\left|g_{13} U-g_{23} U\right| \lesssim \mu R\left(\left\|u-u\left(\cdot+X_{2}-X_{1}\right)\right\|_{\partial B_{1}}+\left\|v_{1}\right\|_{\partial B_{1}}+\left\|v_{2}\right\|_{\partial B_{2}}\right) \tag{3.24}
\end{equation*}
$$

Let us start with $\left\|u-u\left(\cdot+X_{2}-X_{1}\right)\right\|_{\partial B_{1}}$. According to Lemma 3.1 and

## Lemma 3.2 we obtain

$$
\begin{align*}
\| u & -u\left(\cdot+X_{2}-X_{1}\right) \|_{\partial B_{1}} \\
& \lesssim d_{12}\left(\frac{R}{d_{13}^{2}}+\frac{R}{d_{23}^{2}}\right)\left\|-U+\sum_{j \neq 3} A_{3 j} A_{j 3}^{-1} U\right\|_{\partial B_{3}} \\
& +\sum_{i \neq 1,2,3} d_{12}\left(\frac{R}{d_{1 i}^{2}}+\frac{R}{d_{2 i}^{2}}\right)\left\|A_{i 3} A_{33}^{-1} U+\sum_{j \neq 3, i} A_{i j} A_{j 3}^{-1} U\right\|_{\partial B_{i}} \\
& \lesssim d_{12}\left(\frac{R}{d_{13}^{2}}+\frac{R}{d_{23}^{2}}\right)\left(1+\sum_{j \neq 3}\left(\frac{R}{d_{j 3}}\right)^{2}\right)|U| \\
& +\sum_{i \neq 1,2,3} d_{12}\left(\frac{R}{d_{1 i}^{2}}+\frac{R}{d_{2 i}^{2}}\right)\left(\frac{R}{d_{3 i}}+\sum_{j \neq 3, i} \frac{R}{d_{i j}} \frac{R}{d_{3 j}}\right)|U| . \tag{3.25}
\end{align*}
$$

We observe that because of the triangle inequality $d_{3 i} \leq d_{i j}+d_{3 j}$, which we use in form of $\frac{1}{d_{i j} d_{3 j}} \leq \frac{1}{d_{3 i}}\left(\frac{1}{d_{i j}}+\frac{1}{d_{3 j}}\right)$, we have

$$
\begin{equation*}
\sum_{j \neq 3, i} \frac{R}{d_{i j}} \frac{R}{d_{3 j}} \leq \frac{R}{d_{3 i}}\left(\sum_{j \neq i} \frac{R}{d_{i j}}+\sum_{j \neq 3} \frac{R}{d_{3 j}}\right) \lesssim \frac{R}{d_{3 i}} \frac{R N^{2 / 3}}{d_{\min }}=\frac{R}{d_{3 i}} \Lambda \ll \frac{R}{d_{3 i}} \tag{3.26}
\end{equation*}
$$

Furthermore,

$$
\sum_{j \neq 3}\left(\frac{R}{d_{j 3}}\right)^{2} \lesssim \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}=\frac{\Lambda^{2}}{N} \ll 1 .
$$

Hence (3.25) simplifies to

$$
\begin{align*}
& \left\|u-u\left(\cdot+X_{2}-X_{1}\right)\right\|_{\partial B_{1}} \\
& \quad \lesssim d_{12}\left(\left(\frac{R}{d_{13}^{2}}+\frac{R}{d_{23}^{2}}\right)+\sum_{i \neq 1,2,3}\left(\frac{R}{d_{1 i}^{2}}+\frac{R}{d_{2 i}^{2}}\right) \frac{R}{d_{3 i}}\right)|U| . \tag{3.27}
\end{align*}
$$

Using the triangle inequality again, we have

$$
\begin{aligned}
& \sum_{i \neq 1,2,3}\left(\frac{R}{d_{1 i}^{2}}+\frac{R}{d_{2 i}^{2}}\right) \frac{R}{d_{3 i}} \\
& \quad \lesssim\left(\frac{1}{d_{13}}+\frac{1}{d_{23}}\right)\left(\sum_{i \neq 1} \frac{R^{2}}{d_{1 i}^{2}}+\sum_{i \neq 2} \frac{R^{2}}{d_{2 i}^{2}}+\sum_{i \neq 3} \frac{R^{2}}{d_{3 i}^{2}}\right) \\
& \quad \lesssim\left(\frac{1}{d_{13}}+\frac{1}{d_{23}}\right) \frac{R^{2} N^{1 / 3}}{d_{\text {min }}^{2}}
\end{aligned}
$$

Therefore (3.27) turns as desired into

$$
\begin{equation*}
\left\|u-u\left(\cdot+X_{2}-X_{1}\right)\right\|_{\partial B_{1}} \lesssim d_{12}\left(\left(\frac{R}{d_{13}^{2}}+\frac{R}{d_{23}^{2}}\right)+\left(\frac{1}{d_{13}}+\frac{1}{d_{23}}\right) \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}\right)|U| . \tag{3.28}
\end{equation*}
$$

We now turn to $\left\|v_{1}\right\|_{\partial B_{1}}$. According to Lemma 3.1 and Lemma 3.2 we have

$$
\begin{aligned}
\left\|v_{1}\right\|_{\partial B_{1}} & \lesssim \frac{R}{d_{12}}\left\|A_{23} A_{33}^{-1} U+\sum_{j \neq 2,3} A_{2 j} A_{j 3}^{-1} U\right\|_{\partial B_{2}} \\
& \lesssim \frac{R}{d_{12}}\left(\frac{R}{d_{23}}+\sum_{j \neq 2,3} \frac{R}{d_{2 j}} \frac{R}{d_{j 3}}\right)|U| .
\end{aligned}
$$

We now use (3.26) for $i=2$ and so obtain

$$
\begin{equation*}
\left\|v_{1}\right\|_{\partial B_{1}} \lesssim \frac{R}{d_{12}} \frac{R}{d_{23}}|U| . \tag{3.29}
\end{equation*}
$$

Since we trivially have

$$
\frac{R}{d_{12}} \frac{R}{d_{23}} \leq d_{12} \frac{1}{d_{23}} \frac{R^{2}}{d_{\min }^{2}} \leq d_{12} \frac{1}{d_{23}} \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}
$$

also (3.29) turns as desired into

$$
\begin{equation*}
\left\|v_{1}\right\|_{\partial B_{1}} \lesssim d_{12} \frac{1}{d_{23}} \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}|U| . \tag{3.30}
\end{equation*}
$$

By symmetry, we likewise have

$$
\begin{equation*}
\left\|v_{2}\right\|_{\partial B_{2}} \lesssim d_{12} \frac{1}{d_{13}} \frac{R^{2} N^{1 / 3}}{d_{\min }^{2}}|U| . \tag{3.31}
\end{equation*}
$$

Combining (3.24), (3.28), (3.30) \& (3.31), we obtain (1.12).

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